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Bošković's determination of the solar rotation elements in the 18th century using his own observations of sunspot's positions

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Figure 1: Sunspots on the disk of the Sun.

Abstract

Ruder Bošković (1711-1787) observed sunspots on the disk of the Sun during 12th to 29th September 1777 in France. His method is based on: 1) his first dissertation about Sunspots *De Maculis solaribus* (1736) in Latin, describing his geometrical and planar trigonometrical methods in 7 pages with 6 pictures, and 2) detailed calculations elaborated in *Opera pertinentia astronomiam ad opticam, Tomus V, Opusculum II* (1785) in old French, recently translated in Croatian, in 104 pages and 9 pictures. Bošković's method, formulas and drawings are described in

detail in the Opusculum II with an example numerically elaborated in detail, containing his own observational data. In the appendix all his observations are presented. In the 8th and 14th chapters Bošković describes his reflexions on science in the view of contemporary scientific research and related to the scientists which were then active. One of the milestones of this work is the translation of the old French text into Croatian language which reveals Bošković's way of thinking about the topic under consideration.

Keywords: Sunspots, Bošković, solar rotation.

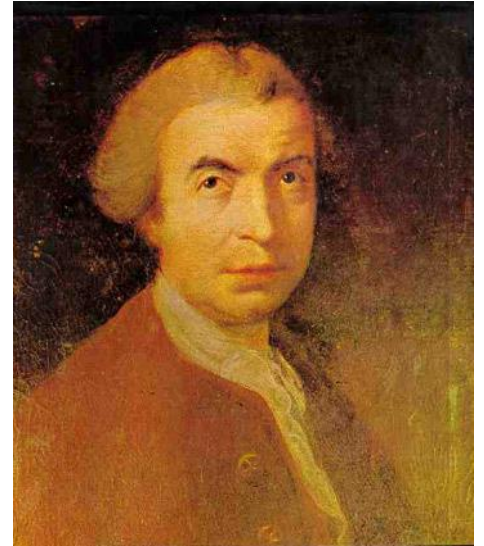
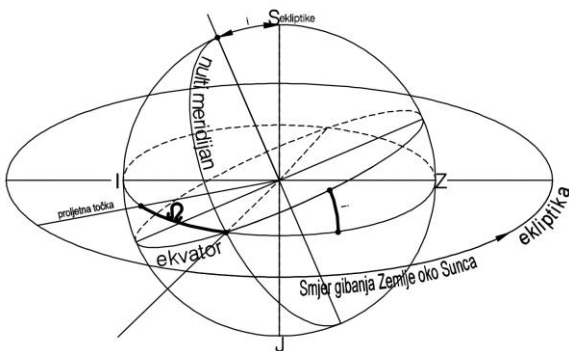


Figure 2: Ruder Bošković (1711-1787).



$$\Omega = 74,40^\circ + (Y - 1850) \cdot 0,014^\circ$$

$$i = 7,25^\circ$$

Figure 3: Carrington's solar rotation elements i and Ω .

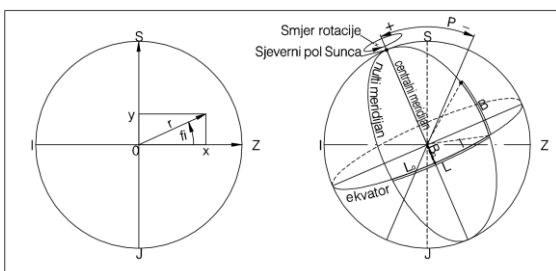
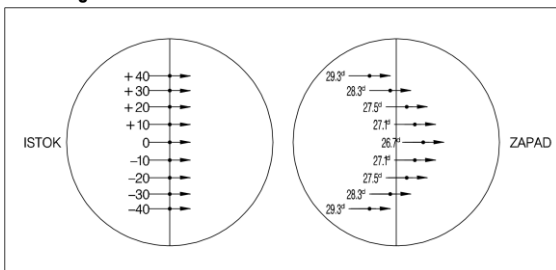


Figure 4: Solar differential rotation (up) and heliographic coordinates (down).

OPUSCULUM II. 89

19. Sept.

bord boréal... 239 ; 241 ; 240 ; 240 ; 240 : milieu 240

1 bord .. 2 ^h . 34 ^m . 26 ^s	2 ^h . 37 ^m . 14 ^s	2 ^h . 41 ^m . 51 ^s	2 ^h . 43 ^m . 47 ^s	2 ^h . 46 ^m . 43 ^s
tache .. 2 . 35 . 3	2 . 37 . 50	2 . 42 . 26	2 . 44 . 23	2 . 47 . 19
2 bord .. 2 . 36 . 34	2 . 39 . 22	2 . 43 . 59	2 . 45 . 55	2 . 48 . 51

Différence ... - 27^{''} ; - 28^{''} ; - 29^{''} ; - 28^{''} ; - 28^{''} : milieu - 28^{''}

Figure 5: Typical set of observations for a day (Bošković, R. J. 1785, 89).

Tab. IX.	Tab. X.	Tab. XI.
4: 1 26 ^m . 6 ^s	4: 1 26 ^m . 6 ^s	A = 365 ^d . 25 . . . 562500
5: 1 26 ^m . 7 ^s	5: 1 26 ^m . 7 ^s	T ¹ = 26,77 . . . 1,427648
6: 1 26 ^m . 6 ^s	6: 1 26 ^m . 6 ^s	(A - T ¹) = 338,48 . . . 7,470458
T 26 ^m . 7 ^s . . . 1,5085426	5: 2 27 ^m . 04	T ¹ = 28,89 . . . 1,460706
M = 54 ^d . 21 ^h . = 3261 . . . 6,486649	6: 2 26 ^m . 82	
9000 2,5934242	6: 3 26 ^m . 67	
T ¹ = 26 ^m . 69 1,426317	26 ^m . 62	
	26 ^m . 77	

Figure 6: Sidereal rotation period of the Sun was determined to be 26.77 days, which is different from the previous value of 25.5 days (Bošković, R. J. 1785, 168, Tab. IX., Tab. X. and Tab. XI.).

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